## ASTR 504B COURSE OUTLINE Winter 2012

✓ The Sun's exterior

<u>Basic properties</u> Distance, radius, mass, density, surface gravity

<u>Photosphere</u> – the visible surface of the Sun Limb darkening (and limb brightening) Granulation Differential rotation – outside and in Sunspots The Sun's magnetic field – Babcock-Leighton model Solar activity cycles

**Chromosphere** 

Corona and solar wind

Flux and luminosity – the solar constant

✓ Principles of stellar structure

Starting assumptions and simplifications <u>Equations of stellar structure</u> Hydrostatic equilibrium Dynamical timescale and other timescales The Virial Theorem Estimating basic internal properties of the Sun Radiation pressure <u>Energy generation</u> – source of the Sun's luminosity <u>Energy transport</u> – How does that energy escape? <u>Convection</u> Radiation and opacity <u>Stellar spectra</u> and absorption lines Mean molecular weight <u>Equations of state</u> Ideal gas law Kinetic theory of gases Degenerate gas Radiation pressure <u>Making models</u> Homologous models Polytropes

- ✓ The Sun's interior (Standard Solar Model)
- ✓ Helioseismology
  - *p- and g-modes Lamb and Brunt–Väisälä frequencies Mode trapping and acoustic cutoff Harmonic oscillations in 3D*

2D Fourier transforms – spatial and temporal Mathematical inversion of the data Kernels and eigenfunctions

*Testing the Standard Solar Model Internal solar rotation Time-distance helioseismology* 

- ✓ Properties of stars
- ✓ Stellar models
- Stellar pulsation and asteroseismology
- ✓ Star and planet formation
- ✓ Stellar rotation